095946

Airborne Geophysical Survey Performed from October 6 to 9, 2010

On the Wounded Moose Property AM 1 to 100 YD28701 to YD28800 AM 101 to 280 YD28901 to YD29080

Dawson Mining District, Yukon NTS Sheet 115O10 63°33' N. Lat., 138°39' W. Long.

Operated by and Recorded to



By Jenny Poon, B.Sc., GIT November 2010

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Airborne Geophysical Survey Report

Precision GeoSurveys Inc.

Wounded Moose Property

PUBB

Prepared for: Taku Gold Corp.

November, 2010 Jenny Poon, B.Sc., GIT

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1.0 Introduction:

This report outlines the survey operations and data processing actions taken during the airborne geophysical survey flown at the Wounded Moose Block. The airborne geophysical survey was flown by Precision GeoSurveys Inc. for Taku Gold Corp. The geophysical survey, carried out from October 6, 2010 to October 9, 2010, saw the acquisition of gamma ray spectrometer data and magnetic data.



Figure 1: Wounded Moose Block area location relative to Dawson, YT.

The Wounded Moose Block is located south-east of Dawson, YT and the Indian River (Figure 1). It is located approximately 64.3 km south-east of Dawson, YT (Figure 2). The survey area itself is approximately 16 km by 4 km. A total of 673.3 line kilometers of radiometric and magnetic data were flown for this survey; this total includes tie lines and survey lines. The survey lines were flown at 100 meter spacings at a $050^{\circ}/230^{\circ}$ heading; the tie lines were flown at 1 km spacings at a heading of $140^{\circ}/320^{\circ}$.





Figure 2: Survey and tie lines outlined in yellow and the boundary in red.

2.0 Geophysical Data:

Geophysical data are collected in a variety of ways and are used to aid in the exploration and determination of geology, mineral deposits, oil and gas deposits, contaminated land sites and UXO detection.

For the purposes of this survey, airborne gamma ray spectrometer and magnetic data were collected to serve in the exploration of the Wounded Moose Block which contains rocks that are prospective for gold mineralization.

2.1 <u>Magnetic Data:</u>

Magnetic surveying is probably the most common airborne survey type to be conducted for both mineral and hydrocarbon exploration. The type of survey specifications, instrumentation, and interpretation procedures, depend on the objectives of the survey. Typically magnetic surveys are performed for:



- 1. Geological Mapping to aid in mapping lithology, structure and alteration in both hard rock environments and for mapping basement lithology, structure and alteration in sedimentary basins or for regional tectonic studies.
- 2. Depth to Basement mapping for exploration in sedimentary basins or mineralization associated with the basement surface.

2.2 Radiometric Data:

Radiometric surveys detect and map natural radioactive emanations, called gamma rays, from rocks and soils. All detectable gamma radiation from earth materials come from the natural decay products of three primary elements: uranium, thorium, and potassium. The purpose of radiometric surveys is to determine either the absolute or relative amounts of U, Th, and K in surface rocks and soils.

3.0 <u>Survey Operations:</u>

Precision GeoSurveys flew the Wounded Moose Block using a Bell 206 BIII Jet Ranger (Figure 3). The survey lines were flown at a nominal line spacing of one hundred (100) meters and the tie lines were flown at 1 km spacing for both the spectrometer and magnetometer as they were acquired simultaneously. The average survey elevation was 31 meters vertically above ground. The experience of the pilots helped to ensure that the data quality objectives were met and that the safety of the flight crew was never compromised given the potential risks involved in airborne surveying.



Figure 3: Bell 206 Jet Ranger equipped with mag stinger for magnetic data acquisition.



The base of operations for this survey was the Gimlex camp located approximately 36 km south-east of Dawson, YT. The Precision crew consisted of a total of four members:

Harmen Keyser and Ola Vaage – Pilots Peter Barker – Operator Jenny Poon – Operator/ On-site Geophysicist

The survey was started on October 6, 2010 to October 9, 2010 with variable snow covered conditions. The survey was complete with some delays due to fog and low cloud ceilings.

4.0 Equipment:

For this survey a magnetometer, spectrometer, base station, laser altimeter, and a data acquisition system were required to carry out the survey and collect quality, high resolution data.

4.1 <u>AGIS:</u>

The Airborne Geophysical Information System, AGIS, (Figure 4), is the main computer used in data recording, data synchronizing, displaying real-time QC data for the geophysical operator, and generation of navigation information for the pilot display system.



Figure 4: AGIS installed in the Bell 206.

The AGIS was manufactured by Pico Envirotec; therefore the system uses standardized Pico software and external sources are connected to the system via RS-232 serial communication cables. The AGIS data format is easily converted into Geosoft or ASCII file formats by a supplied conversion program called PEIView. Additional Pico software allows for post survey quality control procedures.



4.2 <u>Spectrometer:</u>

The IRIS, or Integrated Radiometric Information System is a fully integrated, gamma radiation detection system containing two downward facing NaI detecting crystals for a total volume of 8.4 litres (Figure 5). Real time data acquisition, navigation and communication tasks are integrated into a single unit that is installed in the rear of the aircraft as indicated below. Information such as total count, counts of various elements (K, U, Th, etc.), temperature, barometric pressure, atmospheric humidity and survey altitude can all be monitored on the AGIS screen for immediate QC. All the radiometric data are recorded at 1 Hz.



Figure 5: IRIS strapped into the cargo box of the helicopter.

4.3 Magnetometer:

The magnetometer used by Precision GeoSurveys is a Scintrex cesium vapor CS-3 magnetometer. The system was housed in a front mounted "stinger" (Figure 6). The CS-3 is a high sensitivity/low noise magnetometer with automatic hemisphere switching and a wide voltage range, the static noise rating for the unit is +/- 0.01 nT. On the AGIS screen the operator can view the raw magnetic response, the magnetic fourth difference and the survey altitude for immediate QC of the magnetic data. The magnetic data are recorded at 10 Hz. A magnetic compensator is also used to remove noise created by the movement of the helicopter as it pitches, rolls and yaws within the Earth's geomagnetic field.





Figure 6: View of the mag stinger.

4.4 Base Station:

For monitoring and recording of the Earth's diurnal magnetic field variation, Precision GeoSurveys uses a Scintrex proton precession Envi Pro magnetometer as its base station (Figure 7). This is mounted as close to the survey block as possible to give high, accurate magnetic field data. The Envi Pro base station, uses the well proven precession technology to sample at a rate of 0.5 Hz. A GPS is integrated with the system to record real GPS time that is used to correlate with the GPS time collected by the airborne CS-3 magnetometer.



Figure 7: Scintrex Envi Pro proton precession magnetometer.

4.5 Laser Altimeter:

The pilot is provided with terrain guidance and clearance with an Acuity AccuRange Ar3000 laser altimeter (Figure 8). This is attached at the aft end of the magnetometer



boom. The AR3000 sensor is a time-of-flight sensor that measures distance by a rapidlymodulated and collimated laser beam that creates a dot on the target surface. The maximum range of the laser altimeter is 300 m off of natural surfaces with 90% reflectance and 3 km off special reflectors. Within the sensor unit, reflected signal light is collected by the lens and focused onto a photodiode. Through serial communications and analog outputs, the distance data are transmitted and collected by the AGIS at 10 Hz.



Figure 8: Acuity AccuRange AR3000 laser altimeter.

5.0 Data Processing:

After all the data are collected after a survey flight several procedures are undertaken to ensure that the data meet a high standard of quality. All data were processed using Pico Envirotec software and Geosoft Oasis Montaj geophysical processing software.

5.1 Magnetic Processing:

During aeromagnetic surveying noise is introduced to the magnetic data by the aircraft itself, movement in the aircraft (roll, pitch and yaw) and the permanent magnetization of the aircraft parts (engine and other ferric objects) are large contributing factors to this noise. To remove this noise a process called magnetic compensation is implemented. The magnetic compensation process starts with a test flight at the beginning of the survey where the aircraft flies in the four orthogonal headings required for the survey ($041^{\circ}/220^{\circ}$ and $124^{\circ}/309^{\circ}$ in the case of this survey) at an elevation where there is no ground effect in the magnetic data. In each heading roll, pitch and yaw maneuvers are performed by the pilot, these maneuvers provide the data that is required to calculate the necessary parameters for compensating the magnetic data. A computer program called PEIComp is used to create a model for each survey to remove the noise induced by aircraft movement; this model is applied to each survey flight so the data can be further processed.

A magnetic base station is set up before every flight to ensure that diurnal activity is recorded during the survey flights. Precision GeoSurveys uses a Geometrics 858 base station and sampled at 0.1Hz. Base station readings were reviewed at regular intervals to



insure that no data were collected during periods with high diurnal activity (greater than 5 nT per minute). The base station was installed at a magnetically noise-free area, away from metallic items such as steel objects, vehicles, or power lines. The magnetic variations recorded from the stationary base station are removed from the magnetic data recorded in flight to ensure that the anomalies seen are real and not due to solar activity.

A lag correction was applied to the total magnetic field data to compensate for the lag in the recording system as the magnetometer sensor flies 6.45 m ahead of the GPS antenna. Following a lag correction of 1.7 seconds, a low-pass filter equivalent to 1 second was then applied to the lag corrected data.

5.2 Radiometric Processing:

Radiometric data are processed by windowing the full spectrum to create channels for U, K, Th and total count. A lag correction was also applied to the radiometric data as Pico compensator introduces a lag of 1.4 secs into the positional coordinates for the radiometric data. The data are then lightly filtered and corrected for survey altitude at standard temperature and pressure. Background radioactive contributions from the aircraft, cosmic radiation and atmospheric radon must also be removed. Finally the data are corrected by removing spectral overlap; this is done using the stripping ratios that have been calculated for the spectrometer by prior calibration, this breaks the corrected elemental values down into the apparent radioelement concentrations.

5.3 Final Data Format

Abbreviations used in the GDB files are as follows:

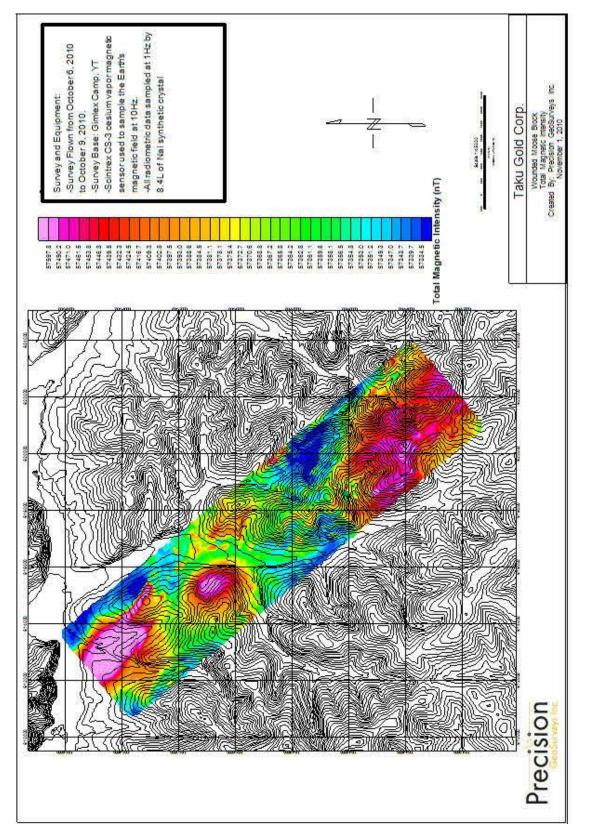
X – Easting in WGS84, UTM zone 7N Y – Northing in WGS84, UTM zone 7N GPStime – GPStime basemag – diurnal data mag – total magnetic field galt – gps altimeter readings lalt – laser altimeter readings dtm – digital terrain model TC_cor – corrected total count K_cor – corrected potassium U_cor – corrected uranium Th_cor – corrected thorium

The file format will be provided in two (2) formats, the first will be a .GDB file for use in Geosoft Oasis Montaj, the second format will be a .XYZ file, this is text file. Two separate files are provided for each format, one for the magnetics and one for the radiometrics.

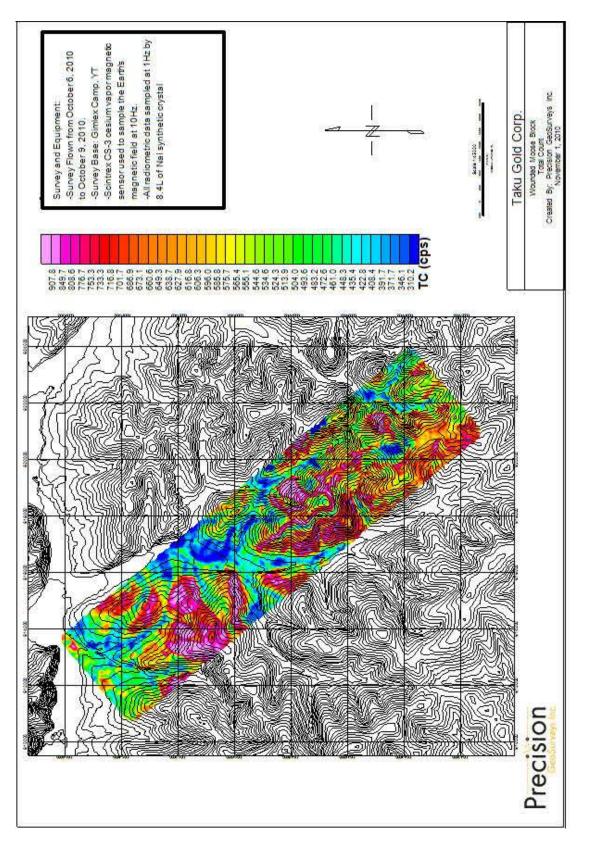


Appendix A Maps

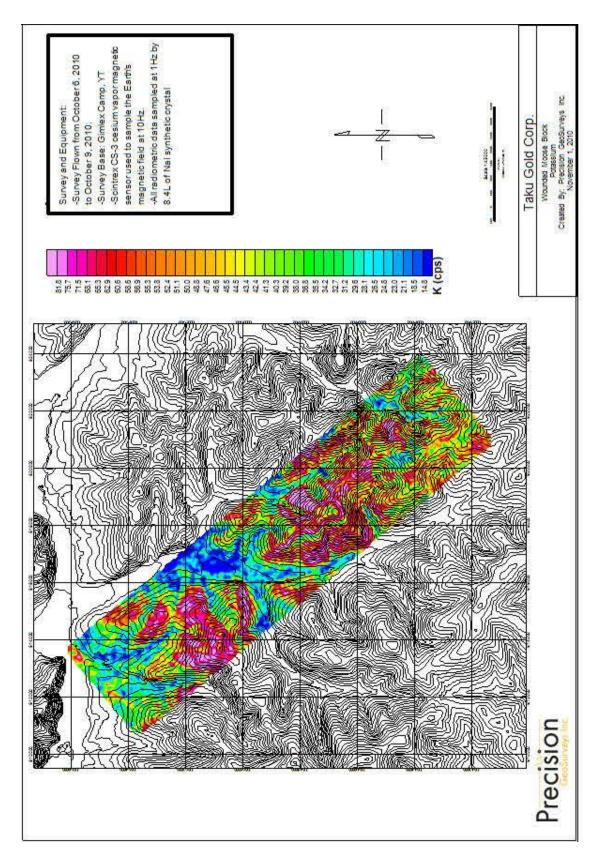






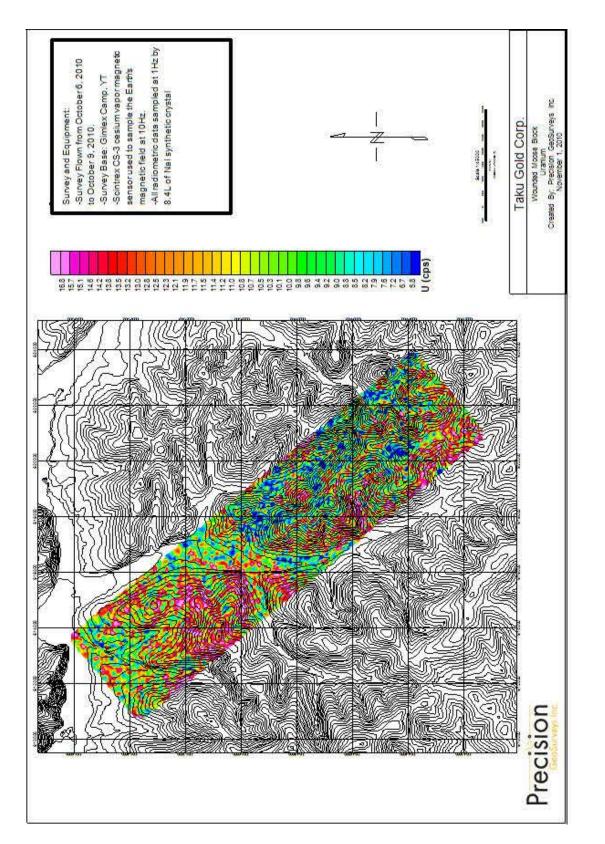




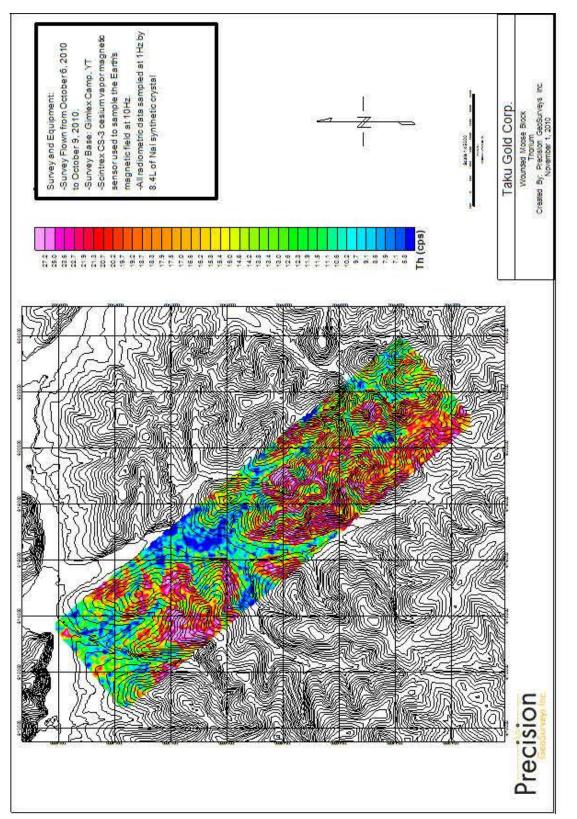






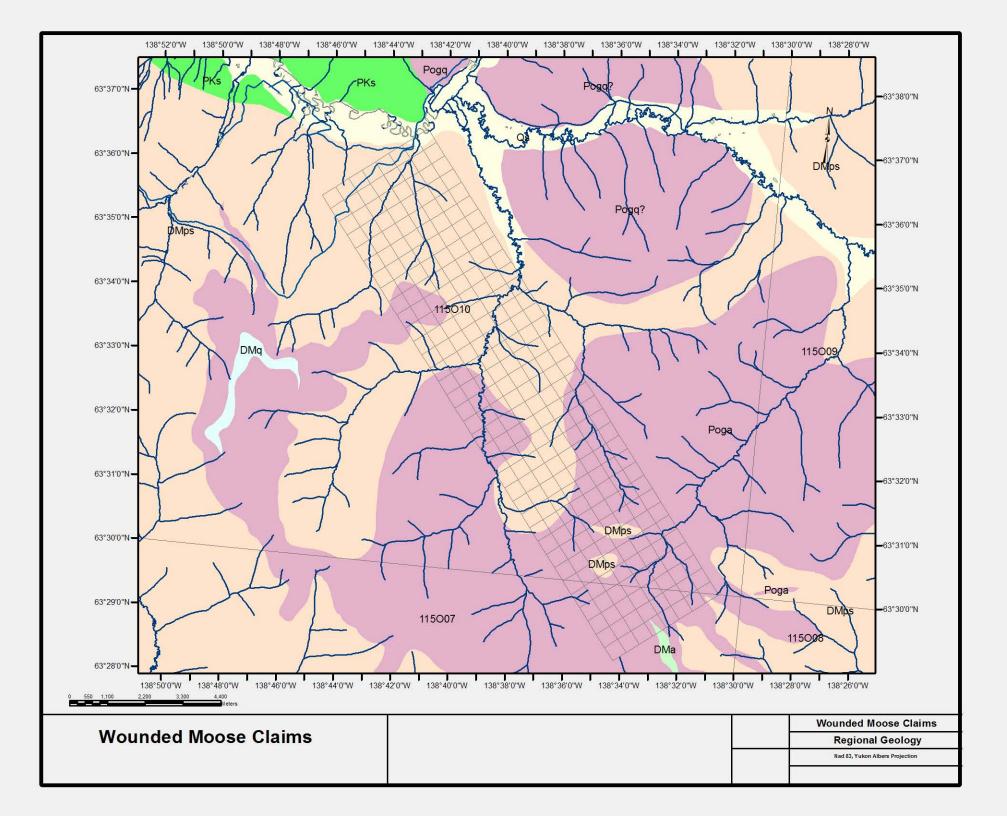






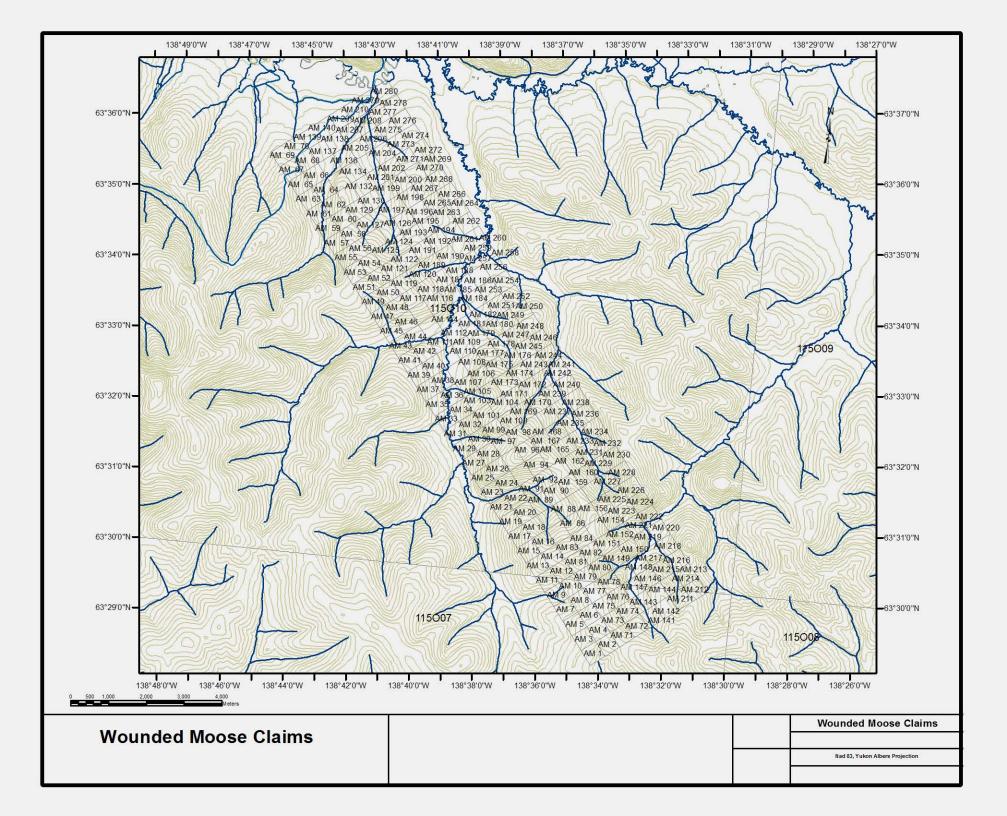






UNIT	AGE	DESCRIPTION							
Qs	Quaternary	Fluvial silt, sand and gravel							
		ORTHOGNEISS (YOUNGER, 264-259 Ma): Pog, undivided orthogneiss; Pogg, pink to orange K-							
		feldspar rich, granitic orthogneiss, commonly includes or associated with Poga; Poga, mainly K-							
Pogq	Permian	feldspar augen orthogneiss, exhibits various states of strain							
		ORTHOGNEISS (YOUNGER, 264-259 Ma): Pog, undivided orthogneiss; Pogg, pink to orange K-							
		feldspar rich, granitic orthogneiss, commonly includes or associated with Poga; Poga, mainly K-							
Poga	Permian	feldspar augen orthogneiss, exhibits various states of strain							
		KLONDIKE SCHIST: muscovite-chlorite-quartz-feldspar schist, chlorite schist, chlorite phyllonite;							
PKs	Permian	local cleaved lapilli tuff with preserved primary texture, probably derived from Pv							
		QUARTZ-MICA SCHIST: undivided metasedimentary rocks dominated by metapsammite,							
		semipelite and metapelite; commonly quartz-garnet-biotite-muscovite schist possibly derived							
DMps	Devonian to Mississippian?	from siliceous siltstone; commonly finely interlayered with garnet metapelite;							
		AMPHIBOLITE: amphibolite schist and gneiss; metabasite; probably derived from mafic to							
		intermediate volcanic or volcaniclastic rocks; locally associated with psammite or interlayered							
DMa	Devonian to Mississippian?	with orthogneiss							
		QUARTZITE: banded to massive, grey to white quartzite; apparantly clastic in origin, or in part,							
DMq	Devonian to Mississippian?	possibly derived from metachert							

Regional Geology Legend (from S.P. Gordey and J.J. Ryan, GSC Open File 4970)



Grant #	Claim		Grant #	Claim		Grant #	Claim		Grant #	Claim	Grant #	Claim
YD28701	AM 1		YD28747	AM 47		YD28907	AM 107		YD28953	AM 153	YD28999	AM 199
YD28702	AM 2		YD28748	AM 48		YD28908	AM 108		YD28954	AM 154	YD29000	AM 200
YD28703	AM 3		YD28749	AM 49		YD28909	AM 109		YD28955	AM 155	YD29001	AM 201
YD28704	AM 4		YD28750	AM 50		YD28910	AM 110		YD28956	AM 156	YD29002	AM 202
YD28705	AM 5		YD28751	AM 51		YD28911	AM 111		YD28957	AM 157	YD29003	AM 203
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YD28713	AM 13		YD28759	AM 59		YD28919	AM 119		YD28965	AM 165	YD29011	AM 211
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YD28716	AM 16		YD28762	AM 62		YD28922	AM 122		YD28968	AM 168	YD29014	AM 214
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YD28718	AM 18		YD28764	AM 64		YD28924	AM 124		YD28970	AM 170	YD29016	AM 216
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YD28726	AM 26		YD28786	AM 86		YD28932	AM 132		YD28978	AM 178	YD29024	AM 224
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